

UG-FT-164/04
CAFPE-34/04**FORMULAE FOR A NUMERICAL COMPUTATION OF
ONE-LOOP TENSOR INTEGRALS^a**ROBERTO PITTAU^b*Departamento de Física Teórica y del Cosmos and Centro Andaluz de Física de
Partículas Elementales (CAFPE), Universidad de Granada, E-18071 Granada,
Spain*

A numerical and iterative approach for computing one-loop tensor integrals is presented.

1 Introduction

In Reference¹ a new approach has been introduced for computing, recursively and numerically, one-loop tensor integrals. Here we describe a few modifications of the original method that allow a more efficient numerical implementation of the algorithm. We keep all of our notations as in Ref.¹ and, in particular, put a bar over n -dimensional quantities and a tilde over ϵ -dimensional objects ($n = 4 + \epsilon$). The formulae we want to modify are Eqs. (9), (35), (48), (52) and (54) of Ref.¹, that, all together, allow to reduce any $(m + 1)$ -point tensor integral with $m \geq 2$ to the standard set of scalar one-loop functions².

Such formulae are not symmetric when interchanging any pair of loop denominators \bar{D}_k , because are derived under the assumption that at least one of them (identified with \bar{D}_0) carries a vanishing external momentum, namely

$$\bar{D}_k = (\bar{q} + p_k)^2 - m_k^2, \quad k = 0, \dots, m, \quad p_0^\mu = 0. \quad (1)$$

Already after the first iteration, terms appear in which the denominator \bar{D}_0 is canceled by a \bar{D}_0 reconstructed in the numerator, so that the resulting integrals do not fulfill any longer the assumption of Eq. (1). A shift of the integration variable \bar{q} is then needed to bring them back to a form suitable to apply the algorithm again. However, shifting \bar{q} may generate a large amount of terms, especially when dealing with high rank tensors, so that deriving more symmetric formulae, in which $p_0^\mu \neq 0$, is clearly preferable.

^aTalk given at LCWS2004, Paris, France, April 2004.

Work supported by the EU under contract HPRN-CT-2002-00149 and by MECN under contract SAB2002-0207.

^bOn leave of absence from Dipartimento di Fisica Teorica, Torino and INFN Sezione di Torino, Italy.

A second useful modification is related to the problem outlined in Sec. 6 of Ref. ¹, that occur when $p_1^2 = 0$ and $p_2^2 \neq 0$ ($p_1^2 \neq 0$ and $p_2^2 = 0$) and $(p_1 \cdot p_2) \sim 0$. For those kinematical configurations a new linear combination of the momenta p_1 and p_2 is needed as a basis of the reduction to ensure the numerical stability of the algorithm. Once again, it is better to include such cases right from the beginning.

2 The General Recursion Formula

When $p_0^\mu \neq 0$, the n -dimensional version of Eq. (9) of Ref. ¹ should be modified as follows

$$\begin{aligned} I_{m; \mu\nu\rho\cdots\tau}^{(n)} &= \frac{\beta}{2\gamma} T_{\mu\nu\lambda\sigma} \left\{ J_{m; \rho\cdots\tau}^{(n)\lambda\sigma} \right\} \\ &- \frac{1}{4\gamma} T_{\mu\nu} \left\{ (m_0^2 - p_0^2) I_{m; \rho\cdots\tau}^{(n)} + I_{m-1; \rho\cdots\tau}^{(n)}(0) - 2p_{0\alpha} I_{m; \rho\cdots\tau}^{(n)\alpha} - I_{m; \rho\cdots\tau}^{(n;2)} \right\} \\ &- \frac{1}{4\gamma} T_{\mu\nu\lambda} \left\{ h_3 I_{m; \rho\cdots\tau}^{(n)\lambda} + I_{m-1; \rho\cdots\tau}^{(n)\lambda}(3) - I_{m-1; \rho\cdots\tau}^{(n)\lambda}(0) - \frac{2\beta}{\gamma} k_{3\alpha} J_{m; \rho\cdots\tau}^{(n)\alpha\lambda} \right\}, \end{aligned} \quad (2)$$

where

$$\begin{aligned} J_{m; \rho\cdots\tau}^{(n)\lambda\sigma} &= (h_1 r_2^\lambda + h_2 r_1^\lambda) I_{m; \rho\cdots\tau}^{(n)\sigma} + (r_2^\lambda + \xi_2 r_1^\lambda) I_{m-1; \rho\cdots\tau}^{(n)\sigma}(1) \\ &+ (r_1^\lambda + \xi_1 r_2^\lambda) I_{m-1; \rho\cdots\tau}^{(n)\sigma}(2) - [r_1^\lambda(1 + \xi_2) + r_2^\lambda(1 + \xi_1)] I_{m-1; \rho\cdots\tau}^{(n)\sigma}(0). \end{aligned} \quad (3)$$

and where the extra integrals $I_{m; \rho\cdots\tau}^{(n;2)}$ are defined in Eq. (77) of Ref. ¹.

In the previous Equations, $k_i = p_i - p_0$ and the massless 4-momenta $\ell_{1,2}$ to be used as a basis of the reduction algorithm, as in Eq. (13) of Ref. ¹, are such that

$$s_1 = \ell_1 + \alpha_1 \ell_2, \quad s_2 = \ell_2 + \alpha_2 \ell_1, \quad (4)$$

where $s_{1,2}$ are suitable linear combinations of $k_{1,2}$

$$s_1 = k_1 + \xi_1 k_2, \quad s_2 = k_2 + \xi_2 k_1. \quad (5)$$

By choosing

$$\xi_1 = \frac{1}{2} \text{sign}(k_2^2) \text{sign}(k_1 \cdot k_2) \quad \text{and} \quad \xi_2 = \frac{1}{2} \text{sign}(k_1^2) \text{sign}(k_1 \cdot k_2), \quad (6)$$

the quantity

$$\gamma = \frac{s_1^2 s_2^2}{(s_1 \cdot s_2) \mp \sqrt{\Delta}} \equiv (s_1 \cdot s_2) \pm \sqrt{\Delta}, \quad \Delta = (s_1 \cdot s_2)^2 - s_1^2 s_2^2, \quad (7)$$

defined in Eq. (62) of Ref. ¹ only vanishes when $k_1^2 = k_2^2 = (k_1 \cdot k_2) = 0$, that always corresponds to collinear configurations cut away in physical observables, therefore solving the second problem outlined in the Introduction. The tensors $T_{\mu\nu\lambda\sigma}, T_{\mu\nu\lambda}, T_{\mu\nu}$ and the 4-vectors r_{12} are defined as in Ref. ¹, but in terms of $\ell_{1,2}$ given in Eq. (4) and with the replacement $p_3 \rightarrow k_3$. Finally

$$\begin{aligned} h_1 &= (m_1^2 - p_1^2) + \xi_1 (m_2^2 - p_2^2) - (1 + \xi_1) (m_0^2 - p_0^2), \\ h_2 &= (m_2^2 - p_2^2) + \xi_2 (m_1^2 - p_1^2) - (1 + \xi_2) (m_0^2 - p_0^2), \\ h_3 &= (m_3^2 - p_3^2) - (m_0^2 - p_0^2), \\ \frac{\beta}{\gamma} &= \pm \frac{1}{2\sqrt{\Delta}}. \end{aligned} \quad (8)$$

The derivation of Eq. (2) closely follows the derivation of Eq. (9) of Ref. ¹. For example, choosing $\ell_{1,2}$ as in Eq. (4), the quantity

$$D_\mu = \frac{1}{\beta} [2(q \cdot \ell_1)\ell_{2\mu} + 2(q \cdot \ell_2)\ell_{1\mu}] \quad (9)$$

defined in Eq. (18) of Ref. ¹ can be rewritten as

$$\begin{aligned} D_\mu &= [\bar{D}_1 + \xi_1 \bar{D}_2 - (1 + \xi_1)\bar{D}_0 + h_1] r_{2\mu} \\ &+ [\bar{D}_2 + \xi_2 \bar{D}_1 - (1 + \xi_2)\bar{D}_0 + h_2] r_{1\mu}, \end{aligned} \quad (10)$$

and generates the term $J^{(n)}$ of Eq. (3). Analogously, choosing $b = k_3$ in Eq. (22) of Ref. ¹, generates the first part of the last term of Eq. (2), because, when $p_0^\mu \neq 0$

$$2(q \cdot k_3) = \bar{D}_3 - \bar{D}_0 + h_3. \quad (11)$$

Finally, the equality

$$q^2 = \bar{D}_0 + (m_0^2 - p_0^2) - 2(q \cdot p_0) - \tilde{q}^2, \quad (12)$$

is the origin of the second row of Eq. (2).

3 Three-point Tensors

When $p_0^\mu \neq 0$, rank 2 and rank 3 three-point tensor integrals need a separate treatment. The relevant formulae follow by adapting the theorems in Eqs. (37) and (40) of Ref. ¹ to the case $p_0^\mu \neq 0$:

$$\int d^n \bar{q} \frac{1}{\bar{D}_0 \bar{D}_1 \bar{D}_2} [(q + p_0) \cdot \ell_3]^i = 0,$$

$$\begin{aligned}
\int d^n \bar{q} \frac{1}{\bar{D}_0 \bar{D}_1 \bar{D}_2} [(q + p_0) \cdot \ell_4]^i &= 0, \quad \forall i = 1, 2, 3 \dots \text{ and} \\
\int d^n \bar{q} \frac{1}{\bar{D}_0 \bar{D}_1 \bar{D}_2} [(q + p_0) \cdot \ell_{3,4}]^2 q_\rho &= 0.
\end{aligned} \tag{13}$$

The final results read as follows

$$\begin{aligned}
I_{2; \mu\nu}^{(n)} &= \frac{\beta}{2\gamma} T'_{\mu\nu\lambda\sigma} \left\{ J_2^{(n)\lambda\sigma} \right\} - \frac{1}{4\gamma} t_{\mu\nu} \left\{ (m_0^2 - p_0^2) I_2^{(n)} + I_1^{(n)}(0) \right. \\
&\quad \left. - 2 p_{0\alpha} I_2^{(n)\alpha} - I_2^{(n;2)} \right\} - \frac{1}{4\gamma} T'_{\mu\nu\lambda} \left\{ p_0^\lambda I_2^{(n)} \right\}, \\
I_{2; \mu\nu\rho}^{(n)} &= \frac{\beta}{2\gamma} T'_{\mu\nu\lambda\sigma} \left\{ J_{2;\rho}^{(n)\lambda\sigma} \right\} - \frac{1}{4\gamma} t_{\mu\nu} \left\{ (m_0^2 - p_0^2) I_{2;\rho}^{(n)} + I_{1;\rho}^{(n)}(0) \right. \\
&\quad \left. - 2 p_{0\alpha} I_{2;\rho}^{(n)\alpha} - I_{2;\rho}^{(n;2)} \right\} - \frac{1}{4\gamma} T'_{\mu\nu\lambda} \left\{ -p_0^\lambda I_{2;\rho}^{(n)} - 2 I_{2;\rho}^{(n)\lambda} \right\}. \tag{14}
\end{aligned}$$

$J^{(n)}$ is given in Eq. (3) and

$$\begin{aligned}
t_{\mu\nu} &= \ell_{3\mu} \ell_{4\nu} + \ell_{4\mu} \ell_{3\nu}, \\
T'_{\mu\nu\lambda} &= -\frac{\ell_{3\mu} \ell_{3\nu} \ell_{4\lambda} (p_0 \cdot \ell_4) + \ell_{4\mu} \ell_{4\nu} \ell_{3\lambda} (p_0 \cdot \ell_3)}{\gamma}.
\end{aligned} \tag{15}$$

4 Rank One Tensors

In this section we adapt Eqs. (48), (52) and (54) of Ref. ¹ to the case $p_0^\mu \neq 0$.

4.1 The $m = 2$ case

With $t_{\alpha\mu}$ defined in Eq. (15) we get

$$I_{2;\mu}^{(n)} = \frac{\beta}{\gamma} J_{2;\mu}^{(n)} + \frac{p_0^\alpha}{2\gamma} t_{\alpha\mu} I_2^{(n)}. \tag{16}$$

4.2 The $m = 3$ case

With $T_{\mu\nu\lambda}$ defined as in Eq. (23) of Ref. ¹ we get

$$\begin{aligned}
I_{3;\mu}^{(n)} &= \frac{\beta}{\gamma} J_{3;\mu}^{(n)} + \frac{1}{4} \left[\frac{\ell_{3\mu}}{(k_3 \cdot \ell_3)} + \frac{\ell_{4\mu}}{(k_3 \cdot \ell_4)} \right] \\
&\quad \times \left\{ h_3 I_3^{(n)} + I_2^{(n)}(3) - I_2^{(n)}(0) - \frac{2\beta}{\gamma} k_3^\lambda J_{3;\lambda}^{(n)} \right\} \\
&\quad - \frac{1}{4\gamma} T_{\mu\nu\lambda} (p_0^\nu k_3^\lambda - p_0^\lambda k_3^\nu) I_3^{(n)}.
\end{aligned} \tag{17}$$

4.3 The $m > 3$ case

The generalization of Eq. (54) of Ref. ¹ reads

$$\begin{aligned}
I_{m;\mu}^{(n)} &= \frac{\beta}{\gamma} J_{m;\mu}^{(n)} + \frac{\ell_{3\mu}\ell_{4\alpha} - \ell_{3\alpha}\ell_{4\mu}}{2\delta} \\
&\times \left\{ k_3^\alpha \left[h_4 I_m^{(n)} + I_{m-1}^{(n)}(4) - I_{m-1}^{(n)}(0) - \frac{2\beta}{\gamma} k_{4\lambda} J_m^{(n)\lambda} \right] \right. \\
&\quad \left. - k_4^\alpha \left[h_3 I_m^{(n)} + I_{m-1}^{(n)}(3) - I_{m-1}^{(n)}(0) - \frac{2\beta}{\gamma} k_{3\lambda} J_m^{(n)\lambda} \right] \right\}, \quad (18)
\end{aligned}$$

where $\delta = (\ell_3 \cdot k_4)(\ell_4 \cdot k_3) - (\ell_3 \cdot k_3)(\ell_4 \cdot k_4)$, and $h_4 = (m_4^2 - p_4^2) - (m_0^2 - p_0^2)$.

5 The Extra Integrals

In most practical cases, the extra integrals, such as $I_{m;\rho\cdots\tau}^{(n;2)}$ in Eq. (2), are either zero or scaleless, so that, even when $p_0^\mu \neq 0$, one can directly use the results given in Appendix B of Ref. ¹. In all other cases modifications are needed. For example, Eqs. (78) and (83) of Ref. ¹ must be replaced by

$$\begin{aligned}
I_{2;\mu}^{(n;2)} &= \frac{i\pi^2}{6} (p_{0\mu} + p_{1\mu} + p_{2\mu}) + \mathcal{O}(\epsilon), \\
I_1^{(n;2)} &= -i\frac{\pi^2}{2} \left[m_1^2 + m_0^2 - \frac{(p_1 - p_0)^2}{3} \right] + \mathcal{O}(\epsilon). \quad (19)
\end{aligned}$$

6 Conclusion

We have derived a set of formulae to efficiently implement the n -dimensional reduction algorithm presented in Ref. ¹. As for the three-point tensors, we limited our analysis to ranks ≤ 3 . For higher ranks, a general formula can be easily derived, with the help of Appendix C of Ref. ¹, using the theorems of Eq. (13).

References

1. F. del Aguila and R. Pittau, “Recursive numerical calculus of one-loop tensor integrals,” arXiv:hep-ph/0404120.
2. G. 't Hooft and M. J. G. Veltman, Nucl. Phys. B **153** (1979) 365.